

The Dirac CP Phase δ_D @ T2(H)K & μ DAR

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2014-5-17

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Status

- March 8, 2012, discovery from **Daya Bay** [later from **Reno**]

$$\sin^2 2\theta_r \sim 0.1$$

by 5.2σ & reaching 7.7σ by October, 2012.

$$\begin{aligned}\theta_r &\equiv \theta_{13}, & \theta_s &\equiv \theta_{12}, & \theta_a &\equiv \theta_{23}, \\ \delta m_a^2 &\equiv \delta m_{31}^2, & \delta m_s^2 &\equiv \delta m_{21}^2.\end{aligned}$$

- Unknowns**

- **Hierarchy:**

$$\delta m_a^2 \gtrsim 0?$$

- **Octant:**

$$\theta_a \gtrsim 45^\circ?$$

- **CP:**

$$\delta_D = ?$$

Inspired by the discovery of relatively large θ_r , we studied the physics potential of two types of future neutrino experiments:

- **Reactor anti-neutrino: MH** [1210.8141]
 - medium baseline $50\text{km} < L < 60\text{km}$: **DayaBayII** + **Reno50**
 - lower energy $1\text{MeV} < E_\nu < 8\text{MeV}$
- **Accelerator neutrino: δ_D , MH, θ_a** [split scheme 1308.6522]
 - long baselines: **T2K(T2HK)** (295km) + **Tokai-to-Oki** (653km) + **NO ν A** (812km) + **LBNE** (1300km) + **LBNO** (2300km)
 - medium energy $0.5\text{GeV} < E_\nu < 5\text{GeV}$
- **Atmospheric neutrino: MH, θ_a** [1309.3176, 1312.0457]
 - long baseline $0\text{km} < L < 13000\text{km}$: **SK(HK)** + **INO** + **PINGU** (*)
 - high energy $1\text{GeV} < E_\nu < 20\text{GeV}$
- **μ DAR: δ_D** [1405.xxxx]

The Dirac CP Phase δ_D @ T2(H)K

- To leading order in $\alpha = \frac{\delta M_{21}^2}{|\delta M_{31}^2|} \sim 3\%$, the oscillation probability relevant to measuring δ_D @ T2(H)K,

$$P_{\nu_\mu \rightarrow \nu_e} \approx 4s_a^2 c_r^2 s_r^2 \sin^2 \phi_{31} \\ - 8c_a s_a c_r^2 s_r c_s s_s \sin \phi_{21} \sin \phi_{31} [\cos \delta_D \cos \phi_{31} \pm \sin \delta_D \sin \phi_{31}]$$

for ν & $\bar{\nu}$, respectively. $[\phi_{ij} \equiv \frac{\delta m_{ij}^2 L}{4E_\nu}]$

- $\nu_\mu \rightarrow \nu_\mu$ Exps measure $\sin^2(2\theta_a)$ precisely, but not $\sin^2 \theta_a$.

- Run both ν & $\bar{\nu}$ modes @ first peak $[\phi_{31} = \frac{\pi}{2}, \phi_{21} = \alpha \frac{\pi}{2}]$,

$$P_{\bar{\nu}_\mu \rightarrow \bar{\nu}_e} + P_{\nu_\mu \rightarrow \nu_e} = 2s_a^2 c_r^2 s_r^2,$$

$$P_{\bar{\nu}_\mu \rightarrow \bar{\nu}_e} - P_{\nu_\mu \rightarrow \nu_e} = \alpha \pi \sin(2\theta_{12}) \sin(2\theta_{13}) \sin(2\theta_{23}) \cos \theta_{13} \sin \delta_D.$$

The Dirac CP Phase δ_D @ T2(H)K

Accelerator experiment, such as T2(H)K, uses off-axis beam to compare ν_e & $\bar{\nu}_e$ appearance @ the oscillation maximum.

- **Disadvantages:**

- Efficiency:

- Proton accelerators produce ν more efficiently than $\bar{\nu}$.
- The $\bar{\nu}$ mode needs more beam time [$T_{\bar{\nu}} : T_{\nu} = 2 : 1$].
- Difficult to reduce the uncertainty.

- Degeneracy:

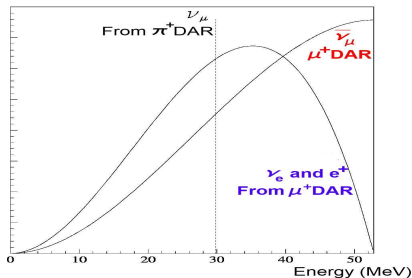
- Only $\sin \delta_D$ appears in $P_{\nu_{\mu} \rightarrow \nu_e}$ & $P_{\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e}$.
- Cannot distinguish δ_D from $\pi - \delta_D$.

- **Solution:**

Measure $\bar{\nu}$ mode with μ^+ decay @ rest (μ DAR)

μ DAR Neutrino Oscillation Experiments

- A cyclotron produces 800 MeV proton beam @ fixed target.
- Produce π^\pm which stop &
 - π^- are absorbed,
 - π^+ decay @ rest: $\pi^+ \rightarrow \mu^+ + \nu_\mu$.
- μ^+ stops & decays @ rest: $\mu^+ \rightarrow e^+ + \nu_e + \bar{\nu}_\mu$.

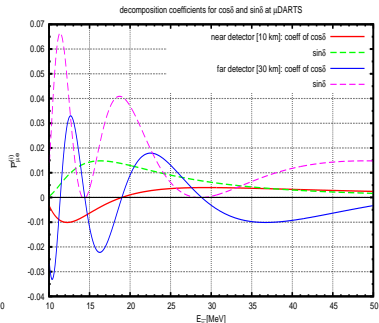
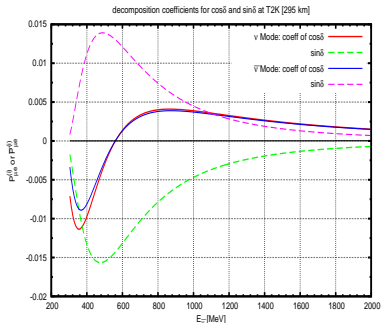


- $\bar{\nu}_\mu$ travel in all directions, oscillating as they go.
- A detector measures the $\bar{\nu}_e$ from $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$.

Accelerator + μ DAR Experiments

Combining $\nu_\mu \rightarrow \nu_e$ @ **accelerator** [narrow peak @ 500 MeV] & $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ @ **μ DAR** [wide peak ~ 35 MeV] solves the 2 problems:

- **Efficiency:**
 - $\bar{\nu}$ @ high intensity μ DAR is plentiful enough.
 - Accelerator Exps can devote all run time to the ν mode. With same run time, the statistical uncertainty drops by $\sqrt{3}$.
- **Degeneracy:** (decomposition in propagation basis [1309.3176])



DAE δ ALUS Project

- It's the **FIRST** proposal along this line:
 - **3** μ DAR with **3** high-intensity cyclotron complexes.
 - **1** detector.
 - Different baselines: **1.5, 8 & 20** km to break degeneracies.
- **Disadvantages:**
 - The scattering lepton from IBD @ low energy is **isotropic**.
 - **Cannot** distinguish $\bar{\nu}_e$ from different sources
 - Cyclotrons **cannot** run simultaneously (20% duty factor).
 - **Large** statistical uncertainty.
 - **Higher intensity** is necessary.
 - **Expensive & Technically challenging**.

New Proposals

1 μ DAR source + 2 detectors

Advantages:

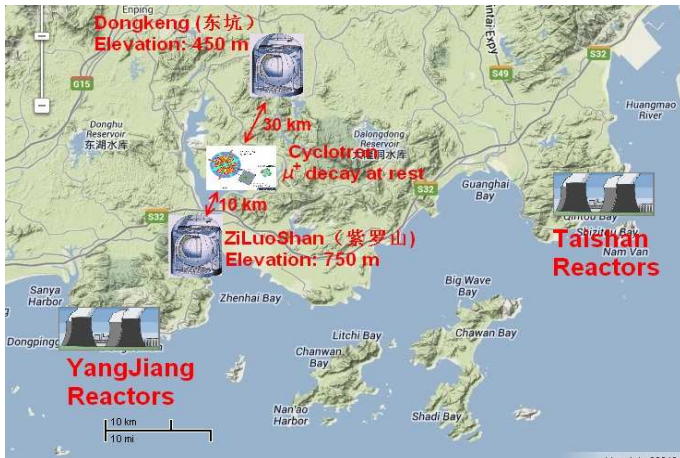
- Full (**100%**) duty factor!
- **Lower** intensity: \sim **9mA** [\sim **4** \times lower than DAE δ ALUS]
- Not far beyond the current state-of-art technology of cyclotron [**2.2mA** @ Paul Scherrer Institute]
- MUCH **cheaper** & technically **easier**.
 - Only one cyclotron.
 - Lower intensity.

Disadvantage:

- A second detector!
 - μ DAR with Two Scintillators (**μ DARTS**) [1401.3977]
 - Tokai 'N Toyama to(2) Kamioka (**TNT2K**) [1405.xxxx]

μ DARTS (1)

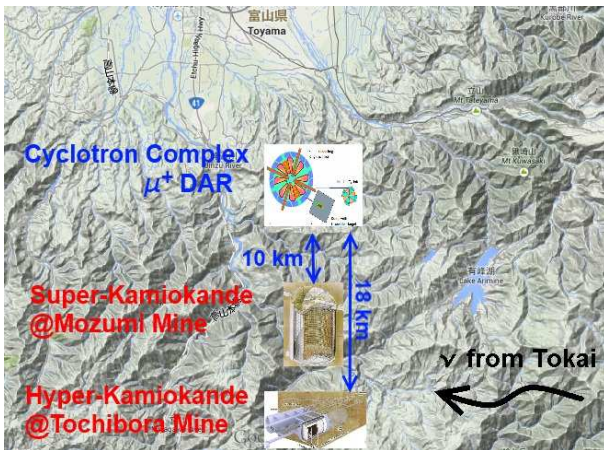
- Two detectors are suggested to overcome the unknown energy response. [Ciuffoli et al., PRD 2014; 1307.7419]



- China Atomic Energy Center has a proposal for cyclotron

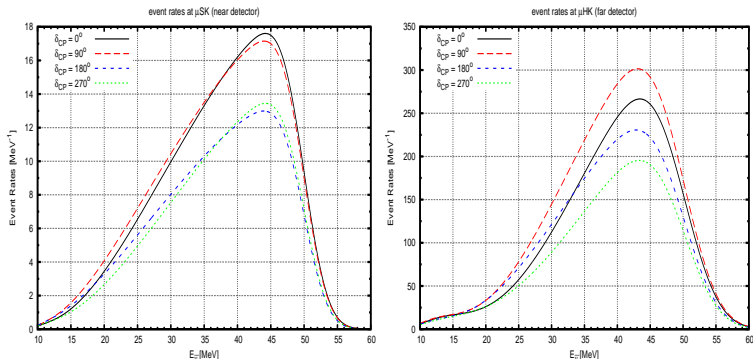
TNT2K (1)

- $TNT2K = T2(H)K + \mu SK + \mu HK$



- μ DAR is also useful for material, medicine sciences

Event Shape @ TNT2K



Expected μ DAR IBD signal from 6 yrs of running @ SK (10km) & HK (18km) with NH.

Backgrounds to IBD ($\bar{\nu}_e + p \rightarrow e^+ + n$)

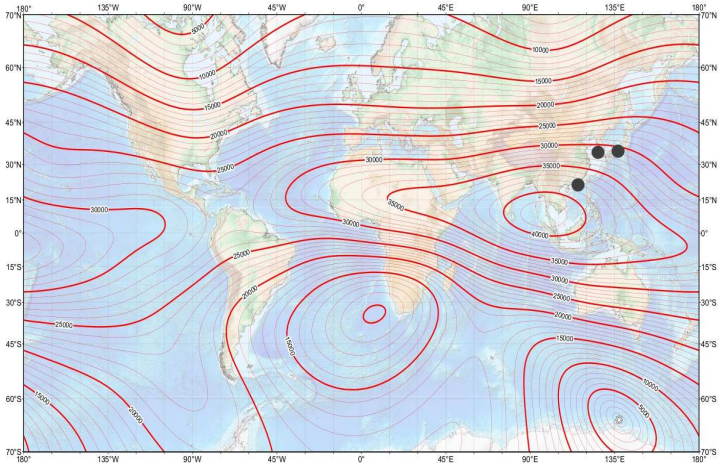
- Reactor $\bar{\nu}_e$: $E_\nu < 10$ MeV
- Accelerator ν_e : $E_\nu > 200$ MeV
- Spallation: $E_\nu \lesssim 20$ MeV
- Supernova Relic Neutrino: $E_\nu \lesssim 20$ MeV

Cut with $30 \text{ MeV} < E_\nu < 55 \text{ MeV}$

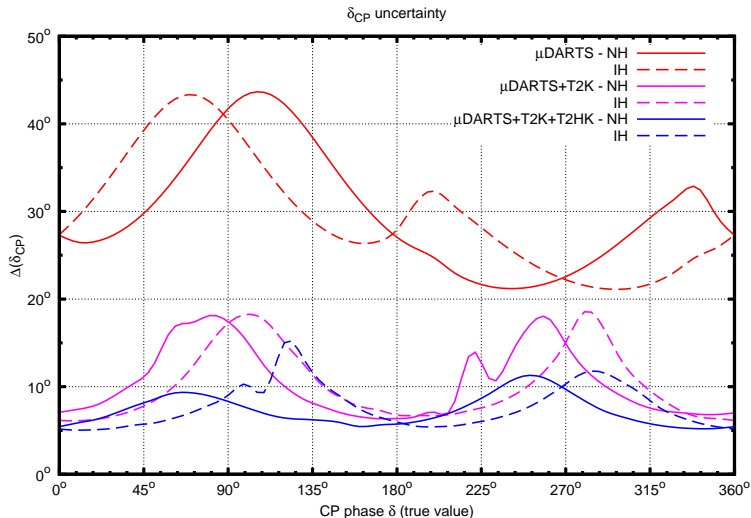
- Accelerator $\nu_\mu \rightarrow$ Invisible muon
- Atmospheric Neutrino Background
 - Invisible muon (below Cherenkov limit)
 - $E_\mu \lesssim 1.5 \times m_\mu$, $\mu^\pm \rightarrow e^\pm$
 - $E_\pi \lesssim 1.5 \times m_\pi$, $\pi^+ \rightarrow \mu^+ \rightarrow e^+$
 - 1 neutron
 - No prompt photon
 - Irreducible $\bar{\nu}_e$: $30 \text{ MeV} \lesssim E_\nu \lesssim 55 \text{ MeV}$
 - Reducible ν_e : $60 \text{ MeV} \lesssim E_\nu \lesssim 100 \text{ MeV}$
 - 1 neutron
 - No prompt photon
 - **Lowest** at μ DARTS & TNT2K sites

Lowest Atmospheric Neutrino Background

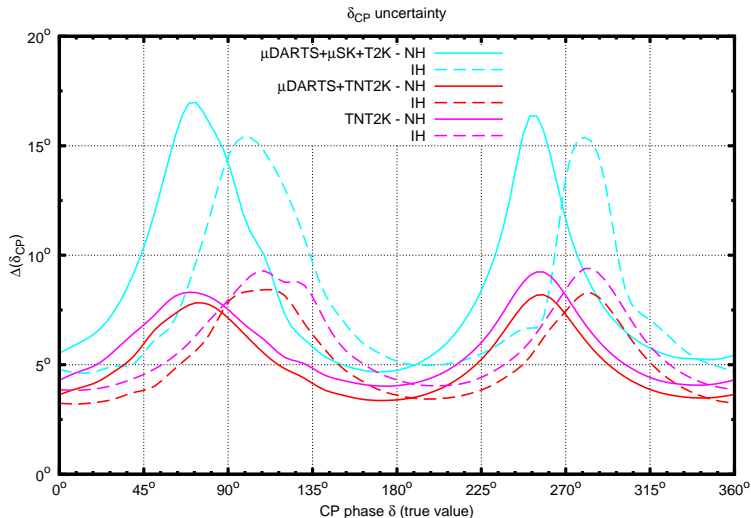
US/UK World Magnetic Model -- Epoch 2010.0
Main Field Horizontal Intensity (H)



δ_D Precision



δ_D Precision



Summary

- μ DAR can run together with accelerator experiments
 - Efficiency
 - Degeneracy
 - Better sensitivity
- 1 μ DAR + 2 detector is much better than DAE δ ALUS
 - Only one μ DAR source
 - Lower intensity
 - Technically easier
 - Much cheaper
- Enhance the physics potential of reactor experiments
- JUNO & T2(H)K site has the lowest atmospheric ν background in the world

Thank You!